

QUANTIFICATION OF MODEL BIAS UNDERLYING THE PHENOMENON OF “EINSTEIN FROM NOISE”

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Abstract: Arising from cryogenic electron microscopy image analysis, “Einstein from noise” refers to spurious patterns that can emerge as a result of averaging a large number of white-noise images aligned to a reference image through rotation and translation. Although this phenomenon is often attributed to model bias, quantitative studies on such bias are lacking. Here, we introduce a simple framework under which an image of p pixels is treated as a vector of dimension p , and a white-noise image is a random vector uniformly sampled from the $(p - 1)$ -dimensional unit sphere. Moreover, we adopt the cross-correlation of two images, which is a similarity measure based on the dot product of image pixels. This framework explains geometrically how the bias results from averaging a properly chosen set of white-noise images that are most highly cross-correlated with the reference image. We quantify the bias in terms of three parameters: the number of white-noise images (n), the image dimension (p), and the size of the selection set (m). Under the conditions that n , p , and m are all large and $(\ln n)^2/p$ and m/n are both small, we show that the bias is approximately $\sqrt{2\gamma/(1 + 2\gamma)}$, where $\gamma = (m/p) \ln(n/m)$.

Key words and phrases: Cross correlation, cryogenic electron microscopy, extreme value distribution, high dimensional data analysis, model bias, white-noise image.

1. Introduction

The phenomenon of “Einstein from noise” comes from the literature on cryogenic electron microscopy (cryo-EM). It refers to an artifact of model bias that arises from averaging a large number of cryo-EM images aligned to a reference (model) image. This artifact of model bias is strongly associated with the noisy nature of cryo-EM images.

Developed for imaging biological macromolecules preserved in a frozen-hydrated state, cryo-EM has become a major tool for the high-resolution structure determination of molecules because of its recent breakthroughs in resolution. In contrast to X-ray crystallography, cryo-EM does not need crystals. Thus, it enables us to determine the structure of proteins that are refractory to crystallization,

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including, in particular, membrane proteins (Liao et al. (2013)) and molecular complexes that exhibit dynamic conformation behaviors (Yan et al. (2015)). In recognition of its success, the Nobel Prize in Chemistry in 2017 was awarded to J. Dubochet, J. Frank, and R. Henderson for their pioneering contributions to the development of cryo-EM.

A technical difficulty encountered by the cryo-EM technique is that the orientations of the molecules are not recorded during the imaging process. As a result, these orientations need to be estimated at the post-imaging stage. However, to mitigate radiation damage, only a minimal dose of electrons can be used to acquire the projection images of individual molecules (called 2D particle images). The resulting cryo-EM images are extremely noisy, with a signal-to-noise ratio less than 0.1. A typical cryo-EM experiment tends to collect a large number of particle images in the hope of increasing the signal-to-noise ratio by suitably averaging the particle images, where the dimension of a particle image is extremely high (larger than 100 by 100). Hence, the data characters of cryo-EM images, including the strong noise contamination, huge dimension, and large sample size, make its processing and statistical analysis very challenging. Henderson (2013) points out how spurious patterns can easily emerge from averaging a large number of white-noise images aligned to a reference image through rotation and translation. He refers specifically to the work of Stewart and Grigorieff (2004), who conducted an experiment by generating 1,000 white-noise images and aligning each of them to Einstein's facial image using rotation and translation. A blurred Einstein's face emerged after averaging the 1,000 aligned images, which Henderson (2013) dubbed "Einstein from noise," showing that an incorrect 3D density map can be constructed if data are blindly fitted to a reference model.

In a recent review paper, Lai et al. (2020) discussed the "Einstein from noise" phenomenon from a statistical perspective. To avoid the technical issue of how rotating an image may destroy the pixel format, they considered a simple mathematical framework under which an image of p pixels is treated as a vector of dimension p , and a white-noise image is a random vector uniformly distributed on the $(p - 1)$ -dimensional unit sphere. The cross-correlation (CC) of two images is adopted, which is a similarity measure based on the dot product of the image pixels and is widely used in image processing. Under this framework, in Section 2, we present a simulation study with $n = 2 \times 10^6$ white-noise images, where the pixel number is $p = 120 \times 120$. Among the 2×10^6 white-noise images, the largest CC value with Einstein's facial image (the reference) is just 0.039. However, the CC increases dramatically to 0.650 after averaging the $m = 800$ images that have the largest CC values with Einstein's facial image. This illustrates the essence

of the “Einstein from noise” phenomenon. This study investigates the “Einstein from noise” phenomenon based on the statistical perspective laid out in Lai et al. (2020). A main task is to approximate the distribution of the CC between the (normalized) average of the m selected images and the reference, which is referred to the (image selection) bias. Although the bias depends on the three parameters n , p , and m in a convoluted manner, when n , p , and m are all large and $(\ln n)^2/p$ and m/n are both small, we show that the bias is approximately $\sqrt{2\gamma/(1+2\gamma)}$, where $\gamma = (m/p) \ln(n/m)$.

The rest of this paper is organized as follows. Section 2 introduces the notation, terminology, and the statistical model and demonstrates the “Einstein from noise” phenomenon. Section 3 consists of two parts: (i) it presents an extreme value theory for the distribution of the largest cross-correlation value as n and p both tend to infinity; and (ii) it states asymptotic results on the bias as n , p , and m all tend to infinity. The theoretical results in part (ii) are validated using a simulation study in Section 4. Section 5 concludes the paper. The proofs of the asymptotic results in Section 3 are relegated to the Appendix. The online Supplementary Material contains the proofs of the auxiliary lemmas.

2. Statistical Model

2.1. Notation, terminology, and model

Let \mathbf{R} be the reference matrix (the digital version of the reference image) of dimension $d_1 \times d_2$. We assume that $\|\mathbf{R}\| = 1$, where $\|\cdot\|$ denotes the Frobenius norm of a matrix or the Euclidean norm of a vector. We generate n independent and identically distributed (i.i.d.) white-noise images, as follows. Let $\mathbf{Z}_1, \dots, \mathbf{Z}_n$ be i.i.d. $d_1 \times d_2$ random matrices, such that the $d_1 d_2$ components of each \mathbf{Z}_i are i.i.d. standard normal. We refer to $\mathbf{Z}_i/\|\mathbf{Z}_i\|$, for $i = 1, \dots, n$ (the normalized version of \mathbf{Z}_i), as n i.i.d. white-noise images.

Let $\mathbf{r} = \text{vec}(\mathbf{R})$, the p -dimensional column vector that is the vectorized version of \mathbf{R} , where $p = d_1 d_2$. The fact that $\|\mathbf{r}\| = 1$ implies that $\mathbf{r} \in \mathcal{S}^{p-1}$ (the $(p-1)$ -dimensional unit sphere). Let $\mathbf{X}_i = \text{vec}(\mathbf{Z}_i)/\|\mathbf{Z}_i\|$. Thus, $\mathbf{X}_1, \dots, \mathbf{X}_n$ are i.i.d. uniformly distributed on \mathcal{S}^{p-1} . We refer to both $\mathbf{Z}_i/\|\mathbf{Z}_i\|$ and \mathbf{X}_i as the i th white-noise image. With \mathbf{r}^\top denoting the transpose of \mathbf{r} , the CC of \mathbf{X}_i and \mathbf{r} (or, equivalently, of $\mathbf{Z}_i/\|\mathbf{Z}_i\|$ and \mathbf{R}) is defined as $\mathbf{r}^\top \mathbf{X}_i$ (the inner product (dot product) of \mathbf{X}_i and \mathbf{r}), which is a similarity measure of two images. Note that $\mathbf{r}^\top \mathbf{X}_i = \cos \Theta_i$, where Θ_i is the angle between \mathbf{r} and \mathbf{X}_i .

The n white-noise images are ordered (and denoted by $\mathbf{X}^{(1)}, \dots, \mathbf{X}^{(n)}$) according to their CC values with \mathbf{r} . In other words, $(\mathbf{X}^{(1)}, \dots, \mathbf{X}^{(n)})$ is a permu-

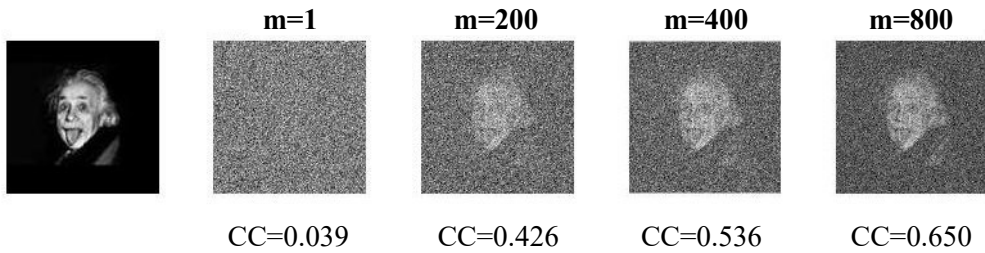


Figure 1. Example with Einstein's face as the reference image.

tation of $(\mathbf{X}_1, \dots, \mathbf{X}_n)$ such that $\mathbf{r}^\top \mathbf{X}^{(1)} \geq \mathbf{r}^\top \mathbf{X}^{(2)} \geq \dots \geq \mathbf{r}^\top \mathbf{X}^{(n)}$. Let $\Theta_{1:n} \leq \Theta_{2:n} \leq \dots \leq \Theta_{n:n}$ be the order statistics of the angles $(\Theta_1, \dots, \Theta_n)$ such that $\cos \Theta_{i:n} = \mathbf{r}^\top \mathbf{X}^{(i)}$, for $i = 1, \dots, n$. Let $\bar{\mathbf{X}}_m = m^{-1} \sum_{i=1}^m \mathbf{X}^{(i)}$. Then, $\bar{\mathbf{X}}_m / \|\bar{\mathbf{X}}_m\| \in \mathcal{S}^{p-1}$ is the normalized average of the m white-noise images that are most highly cross-correlated with the reference image. Our goal is to find a good approximation of the distribution of $\rho_{n,p,m} = \mathbf{r}^\top \bar{\mathbf{X}}_m / \|\bar{\mathbf{X}}_m\|$ when n , p , and m are large. Note that for $m = 1$, $\rho_{n,p,1} = \mathbf{r}^\top \mathbf{X}^{(1)} = \cos \Theta_{1:n}$ is the largest CC value. Note too that the distribution of $\rho_{n,p,m}$ does not depend on \mathbf{r} , because if \mathbf{X} is uniformly distributed on \mathcal{S}^{p-1} , then the distribution of $\mathbf{r}^\top \mathbf{X}$ is independent of \mathbf{r} .

2.2. Demonstration of the “Einstein from noise” phenomenon

We now present two figures summarizing the simulation study described in Section 1, where $n = 2 \times 10^6$, $p = d_1 \times d_2 = 120 \times 120 = 14,400$, and $m = 1, 200, 400, 800$. In Figure 1, the leftmost (reference) image is Einstein's face, and the other four images correspond to $\bar{\mathbf{X}}_m / \|\bar{\mathbf{X}}_m\|$, for $m = 1, 200, 400, 800$. The second image from the left corresponds to $\mathbf{X}^{(1)}$, which has a CC value with Einstein's facial image of 0.039 (the largest among the 2×10^6 white-noise images generated in the simulation). Although this image is rather noisy, Einstein's face emerges in the other three images with different degrees of blurring, corresponding to CC values 0.426, 0.536, and 0.650.

Figure 2 shows similar results based on reference images of a simple chessboard, the digits of 2020, a leopard cat, and the Statistics Building of Academia Sinica, indicating that the phenomenon of “Einstein from noise” is robust across various reference images. The CC values in Figure 2 are roughly the same across the different reference images, which can be explained by the previously mentioned fact that if \mathbf{X} is uniformly distributed on \mathcal{S}^{p-1} , then the distribution of $\mathbf{r}^\top \mathbf{X}$ is independent of \mathbf{r} .

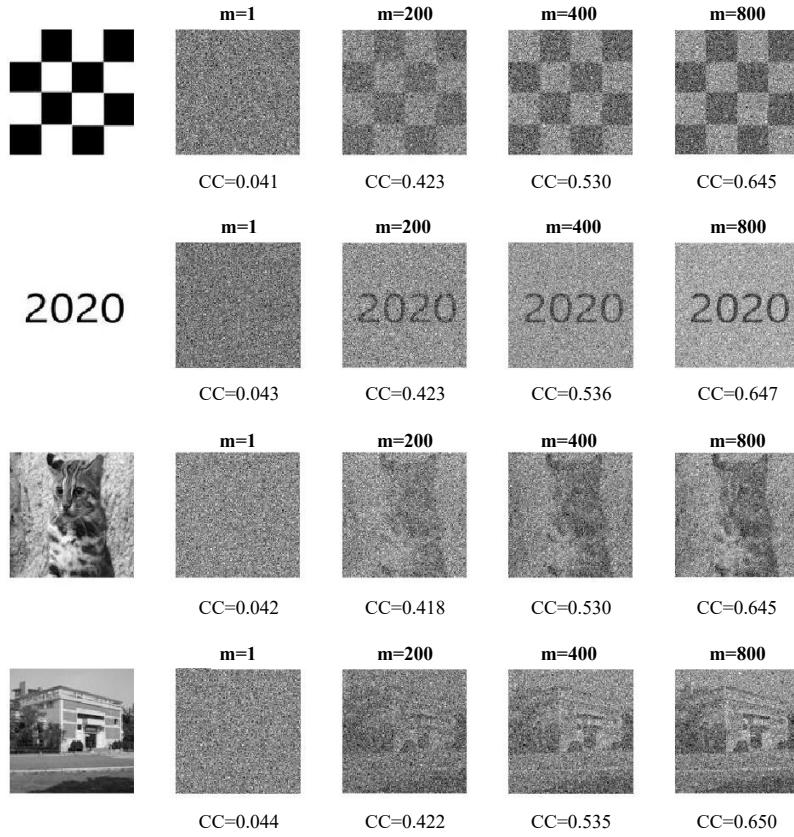


Figure 2. The phenomenon of “Einstein from noise” across various reference images.

3. Asymptotic Theory

3.1. Extreme value theory for the largest CC

Recall that $\cos \Theta_{1:n}$ is the largest CC. The following theorem provides an approximation for the distribution of $\cos \Theta_{1:n}$ when n and p are large.

Theorem 1. *Let*

$$K_{n,p} = -\ln n + \frac{1}{2} \ln \ln n - \frac{1}{2} \ln \left(\frac{2(\ln n)/p}{1 - \exp[(-2 \ln n)/p]} \right) + \frac{1}{2} \ln(4\pi). \quad (3.1)$$

We have

$$(p-1) \ln(\sin \Theta_{1:n}) - K_{n,p} \xrightarrow{d} G \text{ uniformly as } n \wedge p \rightarrow \infty, \quad (3.2)$$

where $n \wedge p = \min\{n, p\}$, \xrightarrow{d} denotes convergence in distribution, and the cumula-

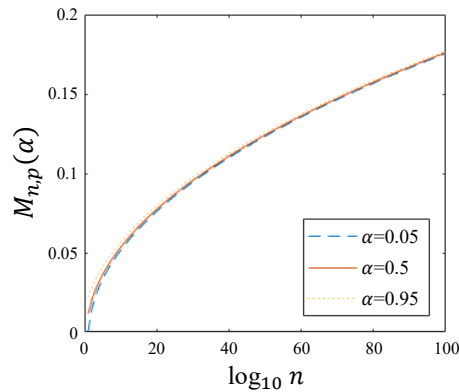


Figure 3. The approximate 100α th quantile of the distribution of $\cos \Theta_{1:n}$ ($M_{n,p}(\alpha)$) versus $\log_{10} n$, with $p = 120 \times 120$, $\alpha = 0.05, 0.5, 0.95$.

tive distribution function (cdf) of G is given by $G(t) = 1 - e^{-e^t}$, for $t \in \mathbb{R}$, which is known as the extreme value distribution of Gumbel type.

Based on (3.2), for $0 < \alpha < 1$, the approximate 100α th quantile of the distribution of $\cos \Theta_{1:n}$ is

$$M_{n,p}(\alpha) = \sqrt{1 - \exp \left\{ \frac{2(K_{n,p} + \ln \ln \alpha^{-1})}{(p-1)} \right\}}.$$

Recall that $\cos \Theta_{1:n} = 0.039$ in the simulation study summarized in Figure 1, where $n = 2 \times 10^6$ and $p = 120 \times 120$. This observed value is compatible with the approximate 10th quantile $M_{n,p}(0.1) = 0.039$.

Figure 3 plots $M_{n,p}(\alpha)$ versus $\log_{10} n$ for $n \leq 10^{100}$, with $p = 120 \times 120$ and $\alpha = 0.05, 0.5, 0.95$. Note that the three quantile curves are very close to each other, indicating that $\cos \Theta_{1:n}$ has a small standard deviation (s.d.). Figure 3 suggests that for $P(\cos \Theta_{1:n} \geq 0.1)$ to be at least 0.05, n is required to be greater than 10^{30} , and for $P(\cos \Theta_{1:n} \geq 0.15)$ to be at least 0.05, n is required to be greater than 10^{70} . In other words, it is unlikely for any of the n i.i.d. white-noise images of dimension 120×120 to have a CC value with Einstein's face greater than 0.15, unless n is astronomically large.

3.2. Asymptotic results on $\rho_{n,p,m}$

When $p = p_n$ and $m = m_n$ both grow with n , the asymptotic expansions for the distribution of $\rho_{n,p,m}$ are more involved. Our analysis requires the condition $(\ln n)^2/p = o(1)$ (which is stronger than $(\ln n)/p = o(1)$), so that terms such as

$(\ln n)(\ln \ln n)/p$ become negligible. Let

$$\beta_{n,p,m} = \frac{m}{p} \left\{ 2 \ln \frac{n}{m} - \ln \ln \frac{n}{m} - \ln(4\pi) + 2 \right\},$$

which is a model bias index. Although the quantity $\beta_{n,p,m}$ plays an important role in our asymptotic results below, we are unaware of any heuristic interpretation of this quantity.

Theorem 2. *Let $p = p_n \rightarrow \infty$ satisfy $(\ln n)^2/p = o(1)$ and $m = m_n \rightarrow \infty$ satisfy $m/n = o(1)$. Then,*

$$\rho_{n,p,m}^2 = \frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}} (1 + o_p(1)).$$

Consequently, $\rho_{n,p,m}^2 - \beta_{n,p,m}/(1 + \beta_{n,p,m}) \rightarrow 0$ in probability.

Theorem 3. *Let $p = p_n \rightarrow \infty$ satisfy $(\ln n)^2/p = o(1)$ and $m = m_n \rightarrow \infty$ satisfy $m(\ln \ln n)^4/(\ln n)^2 = o(1)$. Then,*

$$\alpha_{n,p,m} \left(\rho_{n,p,m}^2 - \frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}} \right) \xrightarrow{d} N(0, 1),$$

where $\alpha_{n,p,m} = p (8m + 2p\beta_{n,p,m}^2)^{-1/2} (1 + \beta_{n,p,m})^2$ and $N(0, 1)$ denotes the standard normal distribution.

Corollary 1. *Let $p = p_n \rightarrow \infty$ and $m = m_n \rightarrow \infty$.*

(i) *If $(\ln n)^2/p = o(1)$ and $m/n = o(1)$, then*

$$\frac{\rho_{n,p,m}}{\sqrt{\beta_{n,p,m}/(1 + \beta_{n,p,m})}} = 1 + o_p(1).$$

Consequently,

$$\rho_{n,p,m} = \sqrt{\frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}}} + o_p(1) \quad \text{and} \quad E(\rho_{n,p,m}) = \sqrt{\frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}}} + o(1).$$

(ii) *In addition to the conditions specified in (i), if $m(\ln \ln n)^4/(\ln n)^2 = o(1)$, then*

$$\tilde{\alpha}_{n,p,m} \left(\rho_{n,p,m} - \sqrt{\frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}}} \right) \xrightarrow{d} N(0, 1),$$

where $\tilde{\alpha}_{n,p,m} = 2\alpha_{n,p,m} \sqrt{\beta_{n,p,m}/(1 + \beta_{n,p,m})}$.

Remark 1. In addition to the condition $(\ln n)^2/p = o(1)$, Theorem 2 only requires the mild condition $m/n = o(1)$. Let $\gamma_{n,p,m} = (m/p) \ln(n/m)$. Because

$\beta_{n,p,m} = 2\gamma_{n,p,m}(1 + o(1))$ (i.e., $2\gamma_{n,p,m}$ is the leading term of $\beta_{n,p,m}$), Theorem 2 implies

$$\rho_{n,p,m}^2 = \frac{2\gamma_{n,p,m}}{1 + 2\gamma_{n,p,m}} + o_p(1).$$

Consequently,

$$\rho_{n,p,m} = \sqrt{\frac{2\gamma_{n,p,m}}{1+2\gamma_{n,p,m}}} + o_p(1) \quad \text{and} \quad E(\rho_{n,p,m}) = \sqrt{\frac{2\gamma_{n,p,m}}{1+2\gamma_{n,p,m}}} + o(1). \quad (3.3)$$

Remark 2. To establish the asymptotic normality of $\rho_{n,p,m}^2$ (and $\rho_{n,p,m}$), Theorem 3 (and Corollary 1) requires the stringent condition $m(\ln \ln n)^4/(\ln n)^2 = o(1)$. It is unclear whether asymptotic normality still holds when m grows at a rate faster than $(\ln n)^2/(\ln \ln n)^4$. Note too that under the conditions in Theorem 3, it is not true that $\alpha_{n,p,m}(\rho_{n,p,m}^2 - 2\gamma_{n,p,m}/(1 + 2\gamma_{n,p,m})) \xrightarrow{d} N(0, 1)$. Thus, while $2\gamma_{n,p,m}$ is the leading term of $\beta_{n,p,m}$, the remaining terms also play a non-negligible role in the proof of asymptotic normality.

Remark 3. Fan, Shao and Zhou (2018) developed an asymptotic theory to approximate the distribution of the maximum spurious correlation of a response variable Y with the best m linear combinations of p covariates \mathbf{X} , based on an i.i.d. sample of size n , when \mathbf{X} and Y are independent; see also Fan, Guo and Hao (2012) for related results. In our setting, the quantity $\rho_{n,p,m}$ is the spurious CC of the reference with the normalized average of the m white-noise images that are most highly cross-correlated with the reference. Indeed, with the roles of n and p reversed, $\rho_{n,p,m}$ corresponds to another spurious correlation of the response variable Y with the the average of the m (standardized) covariates in \mathbf{X} that are most highly correlated with Y when the p covariates in \mathbf{X} and Y are all mutually independent.

4. Simulation Results on $\rho_{n,p,m}$

By Corollary 1(i), if m is small compared to n , and $(\ln n)^2$ is small compared to p , then $E(\rho_{n,p,m})$ is expected to be close to $\sqrt{\beta_{n,p,m}/(1 + \beta_{n,p,m})}$, while the s.d. of $\rho_{n,p,m}$ is expected to be small. We conducted a simulation study of the distribution of $\rho_{n,p,m}$ for various combinations of (n, p, m) , with $n = 10^4, 10^5$, $p = 10^4, 4 \times 10^4$, and $m = 100, 200, 400, 600$. The results are reported in Tables 1 and 2, where $E(\rho_{n,p,m})$ and $\text{s.d.}(\rho_{n,p,m})$ are estimated based on 1,000 replications in each case. Although $\sqrt{\beta_{n,p,m}/(1 + \beta_{n,p,m})}$ approximates $E(\rho_{n,p,m})$ well, it slightly overestimates $E(\rho_{n,p,m})$, more notably for $n = 10^4$. Clearly, $E(\rho_{n,p,m})$ increases as n or m increases or p decreases. On the other hand, $\text{s.d.}(\rho_{n,p,m})$ is small (< 0.005)

Table 1. $p = 10^4$.

m	$n = 10^4$				$n = 10^5$			
	100	200	400	600	100	200	400	600
$E(\rho_{n,p,m})$	0.257	0.323	0.395	0.437	0.318	0.408	0.509	0.570
$\sqrt{\frac{\beta_{n,p,m}}{1+\beta_{n,p,m}}}$	0.258	0.325	0.399	0.442	0.319	0.409	0.510	0.571
s.d. $(\rho_{n,p,m})$	0.0043	0.0045	0.0046	0.0048	0.0039	0.0039	0.0040	0.0037
$\tilde{\alpha}_{n,p,m}^{-1}$	0.0051	0.0053	0.0055	0.0057	0.0041	0.0042	0.0040	0.0039
Prob.	0.974	0.967	0.942	0.870	0.967	0.959	0.947	0.953

Table 2. $p = 4 \times 10^4$.

m	$n = 10^4$				$n = 10^5$			
	100	200	400	600	100	200	400	600
$E(\rho_{n,p,m})$	0.132	0.168	0.210	0.236	0.165	0.218	0.283	0.327
$\sqrt{\frac{\beta_{n,p,m}}{1+\beta_{n,p,m}}}$	0.132	0.169	0.212	0.239	0.166	0.219	0.284	0.328
s.d. $(\rho_{n,p,m})$	0.0022	0.0024	0.0026	0.0027	0.0019	0.0020	0.0021	0.0022
$\tilde{\alpha}_{n,p,m}^{-1}$	0.0026	0.0028	0.0031	0.0033	0.0021	0.0022	0.0023	0.0023
Prob.	0.977	0.978	0.946	0.871	0.968	0.967	0.955	0.953

in all cases. In addition, s.d. $(\rho_{n,p,m})$ decreases as n or p increases, and is about the same as m varies from 100 to 600. Also included in Tables 1 and 2 are $\tilde{\alpha}_{n,p,m}^{-1}$ and the empirical probability (denoted as Prob.) that

$$\left| \rho_{n,p,m} - \sqrt{\frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}}} \right| < 1.96 \tilde{\alpha}_{n,p,m}^{-1}.$$

It is clear from the tables that $\tilde{\alpha}_{n,p,m}^{-1}$ approximates s.d. $(\rho_{n,p,m})$ reasonably well in all cases. By Corollary 1(ii), the Prob. value is expected to be close to 0.95 if the normal approximation is accurate. By Theorem 3 and Corollary 1, $\alpha_{n,p,m}(\rho_{n,p,m}^2 - \beta_{n,p,m}/(1 + \beta_{n,p,m}))$ and $\tilde{\alpha}_{n,p,m}(\rho_{n,p,m} - \sqrt{\beta_{n,p,m}/(1 + \beta_{n,p,m})})$ are approximately standard normal under somewhat stringent conditions on the growth rates of m and p as $n \rightarrow \infty$. None of the combinations of (n, p, m) , with $n = 10^4, 10^5, p = 10^4, 4 \times 10^4$ and $m = 100, 200, 400, 600$, seem to satisfy the condition that $m(\ln \ln n)^4/(\ln n)^2$ be small. Nevertheless, the normal approximation appears to be acceptable for $n = 10^5$, but less satisfactory for $n = 10^4$.

To obtain a more complete picture of the quality of the normal approximation in Corollary 1(ii), in Figures 4–7, we plot the empirical cdf of $\tilde{\alpha}_{n,p,m}(\rho_{n,p,m} - \sqrt{\beta_{n,p,m}/(1 + \beta_{n,p,m})})$ (based on 1,000 replications), along with the standard normal cdf for each combination of (n, p, m) . (The value of D_{ks} is the Kolmogorov–

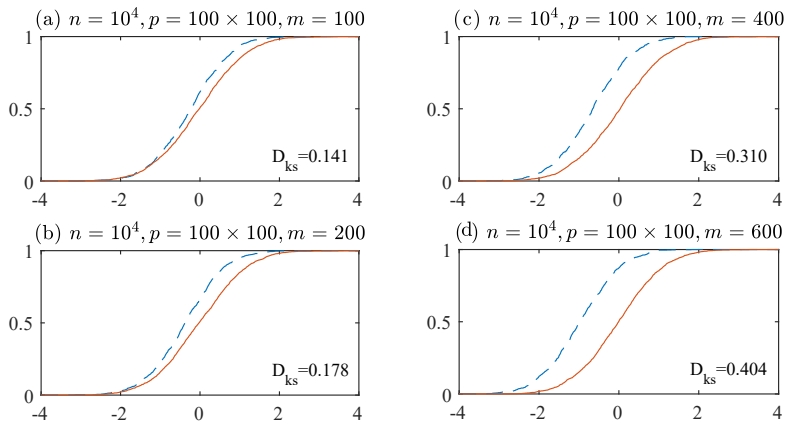


Figure 4. Empirical cdf of $\tilde{\alpha}_{n,p,m}(\rho_{n,p,m} - \sqrt{\beta_{n,p,m}/(1 + \beta_{n,p,m})})$ (dashed curves) and standard normal cdf (solid curves): $n = 10^4$, $p = 10^4$.

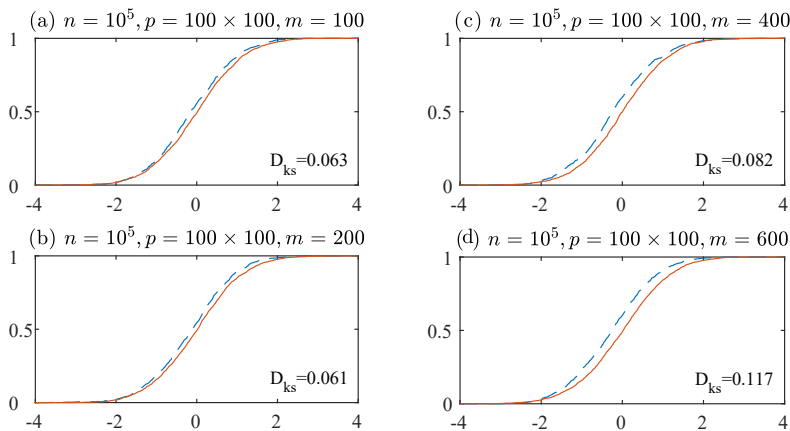


Figure 5. Empirical cdf of $\tilde{\alpha}_{n,p,m}(\rho_{n,p,m} - \sqrt{\beta_{n,p,m}/(1 + \beta_{n,p,m})})$ (dashed curves) and standard normal cdf (solid curves): $n = 10^5$, $p = 10^4$.

Smirnov distance between the two cdfs.) Figures 4–7 show the cdf under four different scenarios, depending on the values of $n = 10^4, 10^5$ and $p = 10^4, 4 \times 10^4$. Each figure includes four plots, depending on the values of $m = 100, 200, 400, 600$. The empirical cdf is shifted to the left of the standard normal cdf (particularly for $n = 10^4$ in Figures 4 and 6), indicating that the mean of $\rho_{n,p,m} - \sqrt{\beta_{n,p,m}/(1 + \beta_{n,p,m})}$ is negative. This is consistent with the results in Tables 1 and 2, where $\sqrt{\beta_{n,p,m}/(1 + \beta_{n,p,m})}$ (slightly) overestimates $E(\rho_{n,p,m})$ (particularly for $n = 10^4$).

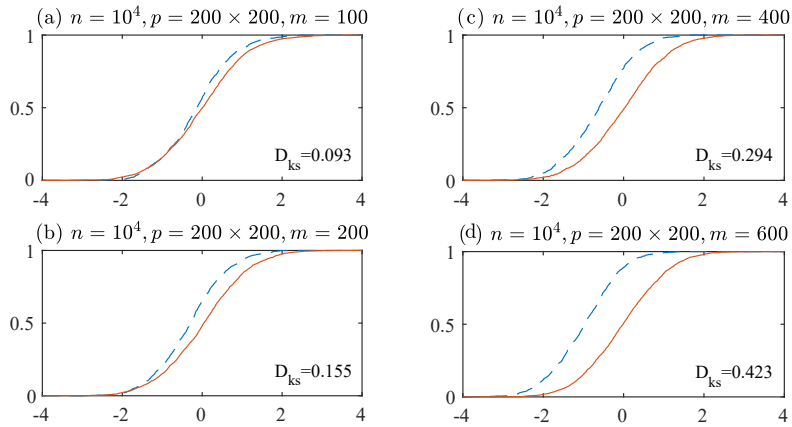


Figure 6. Empirical cdf of $\tilde{\alpha}_{n,p,m}(\rho_{n,p,m} - \sqrt{\beta_{n,p,m}/(1 + \beta_{n,p,m})})$ (dashed curves) and standard normal cdf (solid curves): $n = 10^4$, $p = 4 \times 10^4$.

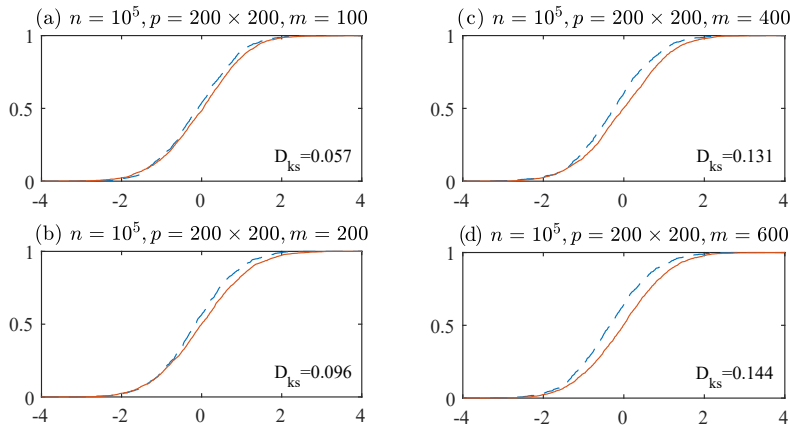


Figure 7. Empirical cdf of $\tilde{\alpha}_{n,p,m}(\rho_{n,p,m} - \sqrt{\beta_{n,p,m}/(1 + \beta_{n,p,m})})$ (dashed curves) and standard normal cdf (solid curves): $n = 10^5$, $p = 4 \times 10^4$.

5. Conclusion

We have studied a simple statistical model in order to quantitatively examine the ‘Einstein from noise’ phenomenon. Specifically, for a given reference image of dimension p and a set S_n of n i.i.d. white-noise images (with the common uniform distribution on \mathcal{S}^{p-1}), we derived the asymptotic behavior of the CC $\rho_{n,p,m}$ between the reference and the normalized average of the m ‘most biased’ members in S_n , in the sense that they have the largest CC values with the reference. Our theoretical results indicate that for $m = 1$ and $p = 120 \times 120$, unless n

is far beyond the practical range ($> 10^{70}$), $\rho_{n,p,1}$ is small (< 0.15) with high probability, implying that none of the n white-noise images even remotely resembles the reference. On the other hand, for m moderately large (≥ 400), $\rho_{n,p,m}$ exceeds 0.5 with high probability if $n = 2 \times 10^6$. In this case, a blurred version of the reference emerges from the normalized average of the m most biased members in S_n .

Given a set S_n of n i.i.d. white-noise images, Cai, Fan and Jiang (2013) derived the asymptotic distribution of the maximum of all pairwise CCs in S_n ; see also Cai and Jiang (2011, 2012), and the references therein. In the absence of a reference image, their results may be applied to test the null hypothesis that S_n consists of n i.i.d. white-noise images. On the other hand, given a reference image, we can use our results to test such a null hypothesis against the alternative that some of the n images in S_n are biased toward the reference by checking whether $\rho_{n,p,m}$ exceeds a threshold (determined by the null distribution of $\rho_{n,p,m}$).

Our approach can be generalized directly to tackle a special case of multiple references. Let $\mathbf{r}^{(1)}, \dots, \mathbf{r}^{(k)}$ be k given references of dimension p . Given a set S_n of n i.i.d. white-noise images, for $i = 1, \dots, k$, let $\rho_{n,p,m}^{(i)}$ ($i = 1, \dots, k$) denote the CC between $\mathbf{r}^{(i)}$ and the normalized average of those m members in S_n with the largest CC values with $\mathbf{r}^{(i)}$. It would be of interest to derive the asymptotic distribution of $\max\{\rho_{n,p,m}^{(i)} : i = 1, \dots, k\}$. If $\mathbf{r}^{(1)}, \dots, \mathbf{r}^{(k)}$ are orthogonal (i.e., the pairwise CCs are all equal to zero), then it can be argued that $\rho_{n,p,m}^{(1)}, \dots, \rho_{n,p,m}^{(k)}$ are asymptotically independent. In this case the asymptotic distribution of $\max\{\rho_{n,p,m}^{(i)} : i = 1, \dots, k\}$ can be readily derived from Corollary 1. However, it seems difficult to find the asymptotic distribution of $\max\{\rho_{n,p,m}^{(i)} : i = 1, \dots, k\}$ when $\mathbf{r}^{(1)}, \dots, \mathbf{r}^{(k)}$ are not orthogonal.

The phenomenon of ‘‘Einstein from noise’’ originally arose in the context of cryo-EM image analysis, where a key component is image alignment (including rotation and translation). While addressing this more complicated problem is beyond the scope of this study, note that the geometric shape of the reference is likely to play a significant role in the asymptotic theory yet to be developed. As an example, consider a rotationally invariant reference, such as an image of a centered wheel. Because of the rotational symmetry of the reference, a data image cannot fit the reference any better after rotation. We leave this challenging problem for future work.

Supplementary Material

The online Supplementary Material contains the proofs of Lemmas A6–A8 in the Appendix.

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A. Appendix

The Appendix consists of three sections. Section A.1 states some auxiliary lemmas, Section A.2 contains the proof of Theorem 1, and Section A.3 provides the proofs of Theorems 2 and 3 and Corollary 1. For easy reference, a complete list of notations is given in Supplementary Material. Note that if \mathbf{X} is uniformly distributed on \mathcal{S}^{p-1} , then the distribution of $\mathbf{r}^\top \mathbf{X}$ is the same for all $\mathbf{r} \in \mathcal{S}^{p-1}$. Without loss of generality, we assume $\mathbf{r} = (1, 0, \dots, 0)^\top \in \mathcal{S}^{p-1}$ in what follows.

A.1. Auxiliary lemmas

Lemma A1. (*Lemma 6.2 of Cai and Jiang (2012)*) For $t \in (0, 1)$, we have

$$\left(1 + \frac{1}{pt^2}\right)^{-1} \frac{1}{(p+2)t} (1-t^2)^{(p+2)/2} \leq \int_t^1 (1-u^2)^{p/2} du \leq \frac{1}{(p+2)t} (1-t^2)^{(p+2)/2}.$$

Since $\mathbf{X}_i, i = 1, \dots, n$ are iid uniformly distributed on \mathcal{S}^{p-1} and Θ_i denotes the angle between \mathbf{X}_i and $\mathbf{r} = (1, 0, \dots, 0)^\top$, we have (cf. Eq (5) of Cai, Fan and Jiang (2013)) that $\Theta_i, i = 1, \dots, n$ are iid with the common cdf

$$\begin{aligned} F_p(\theta) &= \int_0^\theta \frac{1}{\sqrt{\pi}} \frac{\Gamma(p/2)}{\Gamma((p-1)/2)} (\sin x)^{p-2} dx \\ &= \int_{\cos \theta}^1 \frac{1}{\sqrt{\pi}} \frac{\Gamma(p/2)}{\Gamma((p-1)/2)} (1-u^2)^{(p-3)/2} du, \quad \theta \in [0, \pi]. \end{aligned} \tag{A.1}$$

Let

$$\bar{F}_p(\theta) = \frac{1}{\sqrt{\pi}} \frac{\Gamma(p/2)}{\Gamma((p-1)/2)} \frac{\sin^{p-1} \theta}{(p-1)|\cos \theta|}. \tag{A.2}$$

The following lemma is a consequence of Lemma A1.

Lemma A2. For $\theta \in (0, \pi/2)$ and $p > 3$, we have

$$\left(1 + \frac{1}{(p-3)\cos^2\theta}\right)^{-1} \bar{F}_p(\theta) \leq F_p(\theta) \leq \bar{F}_p(\theta).$$

Let U_1, U_2, \dots be iid uniform $(0,1)$ random variables and let $U_{1:n} \leq \dots \leq U_{n:n}$ denote the order statistics of U_1, \dots, U_n . Let $S_0 = 0$, and $S_i = \xi_1 + \dots + \xi_i$, $i = 1, 2, \dots$, where ξ_1, ξ_2, \dots are iid exponential random variables with mean 1. The next lemma is well known; see e.g. Karlin and Taylor (1975). We write $\mathbf{X} \stackrel{d}{=} \mathbf{Y}$ if random vectors \mathbf{X} and \mathbf{Y} are equal in distribution.

Lemma A3. (i) $(U_{1:n}, \dots, U_{n:n}) \stackrel{d}{=} (S_1, \dots, S_n)/S_{n+1}$. (ii) $(S_1, \dots, S_n)/S_{n+1}$ is independent of S_{n+1} .

Recall that $(\mathbf{X}^{(1)}, \dots, \mathbf{X}^{(n)})$ is a permutation of $(\mathbf{X}_1, \dots, \mathbf{X}_n)$ such that $X_1^{(1)} \leq \dots \leq X_1^{(n)}$, where $X_1^{(i)} = \mathbf{r}^\top \mathbf{X}^{(i)}$ (the first component of $\mathbf{X}^{(i)}$). Let \mathbf{V}_i and $\mathbf{V}^{(i)}$ be defined by $\mathbf{X}_i = (X_{i1}, (1 - X_{i1}^2)^{1/2} \mathbf{V}_i^\top)^\top$ and $\mathbf{X}^{(i)} = (X_1^{(i)}, \nu_i \mathbf{V}^{(i)\top})^\top$, where $\nu_i = (1 - X_1^{(i)2})^{1/2}$. In other words, \mathbf{V}_i ($\mathbf{V}^{(i)}$, respectively) $\in \mathcal{S}^{p-2}$ is the normalized subvector of \mathbf{X}_i ($\mathbf{X}^{(i)}$, respectively) with the first component deleted.

Lemma A4.

- (i) X_{i1} and $\mathbf{V}_i, i = 1, \dots, n$ are all independent.
- (ii) $X_{i1}, i = 1, \dots, n$ are iid.
- (iii) $\mathbf{V}_i, i = 1, \dots, n$ are iid with the uniform distribution on \mathcal{S}^{p-2} .
- (iv) $(\mathbf{V}^{(1)}, \dots, \mathbf{V}^{(n)})$ is independent of (X_{11}, \dots, X_{n1}) and hence independent of $(X_1^{(1)}, \dots, X_1^{(n)})$.
- (v) $\mathbf{V}^{(i)}, i = 1, \dots, n$ are iid with the uniform distribution on \mathcal{S}^{p-2} .

To show Lemma A4, let $Z_{ij}, i = 1, \dots, n, j = 1, \dots, p$, be i.i.d. standard normal, and let

$$\mathbf{X}_i^* = (Z_{i1}, \dots, Z_{ip})^\top / \sqrt{\sum_{j=1}^p Z_{ij}^2} = \left(Z_{i1} / \sqrt{\sum_{j=1}^p Z_{ij}^2}, \nu_i^* \mathbf{V}_i^* \right)^\top, \quad i = 1, \dots, n,$$

where $\nu_i^* = \sqrt{\sum_{j=2}^p Z_{ij}^2} / \sqrt{\sum_{j=1}^p Z_{ij}^2}$ and $\mathbf{V}_i^* = (Z_{i2}, \dots, Z_{ip})^\top / \sqrt{\sum_{j=2}^p Z_{ij}^2}$.

It is readily seen that \mathbf{X}_i^* is uniformly distributed on \mathcal{S}^{p-1} and independent of $\sum_{j=1}^p Z_{ij}^2$, and that \mathbf{V}_i^* is uniformly distributed on \mathcal{S}^{p-2} and independent of Z_{i1}

and $\sum_{j=2}^p Z_{ij}^2$ (hence independent of $Z_{i1}/\sqrt{\sum_{j=1}^p Z_{ij}^2}$). Since $(\mathbf{X}_1, \dots, \mathbf{X}_n) \stackrel{d}{=} (\mathbf{X}_1^*, \dots, \mathbf{X}_n^*)$ and $(\mathbf{V}_1, \dots, \mathbf{V}_n) \stackrel{d}{=} (\mathbf{V}_1^*, \dots, \mathbf{V}_n^*)$, Lemma A4 follows.

Recall that

$$\bar{\mathbf{X}}_m = \frac{1}{m} \sum_{i=1}^m \mathbf{X}^{(i)} = \left(m^{-1} \sum_{i=1}^m X_1^{(i)}, m^{-1} \sum_{i=1}^m \nu_i \mathbf{V}^{(i)\top} \right)^\top$$

and that

$$\rho_{n,p,m}^2 = \left(\mathbf{r}^\top \frac{\bar{\mathbf{X}}_m}{\|\bar{\mathbf{X}}_m\|} \right)^2 = \frac{\left((1/m) \sum_{i=1}^m X_1^{(i)} \right)^2}{\left((1/m) \sum_{i=1}^m X_1^{(i)} \right)^2 + \left\| (1/m) \sum_{i=1}^m \nu_i \mathbf{V}^{(i)} \right\|^2}.$$

Let $\mathbf{V}'_i, i = 1, \dots, n$ be iid uniformly distributed on \mathcal{S}^{p-2} and independent of $\mathbf{X}_1, \dots, \mathbf{X}_n$. Then the following lemma is a consequence of Lemma A4.

Lemma A5.

$$\begin{aligned} \rho_{n,p,m}^2 &\stackrel{d}{=} \frac{\left(m^{-1} \sum_{i=1}^m X_1^{(i)} \right)^2}{\left(m^{-1} \sum_{i=1}^m X_1^{(i)} \right)^2 + \left\| m^{-1} \sum_{i=1}^m \nu_i \mathbf{V}'_i \right\|^2} \\ &= \frac{A_{n,p,m}}{A_{n,p,m} + V_{n,p,m}}, \end{aligned} \tag{A.3}$$

where

$$A_{n,p,m} = \left(\frac{1}{m} \sum_{i=1}^m X_1^{(i)} \right)^2 \text{ and } V_{n,p,m} = \left\| \frac{1}{m} \sum_{i=1}^m \nu_i \mathbf{V}'_i \right\|^2. \tag{A.4}$$

The long proofs of Lemmas A6-A8 below are given in Supplementary Material.

Lemma A6. *Let $m = m_n \rightarrow \infty$ satisfy $m/n = o(1)$ and $p = p_n \rightarrow \infty$ satisfy $(\ln n)^2/p = O(1)$. Then*

(i)

$$\max_{1 \leq i \leq m} \left| p \ln(\sin \Theta_{i:n}) + \ln \frac{n}{i} - \frac{1}{2} \ln \ln \frac{n}{i} \right| = O_p(1),$$

(ii)

$$\max_{1 \leq i \leq m} \left| -\frac{p}{2} \cos^2 \Theta_{i:n} + \ln \frac{n}{i} - \frac{1}{2} \ln \ln \frac{n}{i} \right| = O_p(1),$$

where $\Theta_{1:n} \leq \Theta_{2:n} \leq \dots \leq \Theta_{n:n}$ are the order statistics of $\Theta_1, \dots, \Theta_n$.

Lemma A7. *Suppose that $p = p_n \rightarrow \infty$ satisfies $(\ln n)^2/p = O(1)$.*

(i) *If $m = m_n \rightarrow \infty$ satisfies $m/n \rightarrow 0$, then*

$$-pA_{n,p,m} + 2 \ln \frac{n}{m} - \ln \ln \frac{n}{m} = O_p(1).$$

(ii) *If $m = m_n \rightarrow \infty$ satisfies $(\ln m)^3/(\ln n)^2 \rightarrow 0$, then*

$$-pA_{n,p,m} + 2 \ln \frac{n}{m} - \ln \ln \frac{n}{m} - \ln(4\pi) + 2 - \frac{2}{p} \left(\ln \frac{n}{m} \right)^2 = o_p(1).$$

(iii) *If $m = m_n \rightarrow \infty$ satisfies $m(\ln \ln n)^4/(\ln n)^2 \rightarrow 0$, then*

$$\begin{aligned} & \left(\frac{m}{8} \right)^{1/2} \left\{ -pA_{n,p,m} + 2 \ln \frac{n}{m} - \ln \ln \frac{n}{m} - \ln(4\pi) + 2 - \frac{2}{p} \left(\ln \frac{n}{m} \right)^2 \right\} \\ & \xrightarrow{d} N(0, 1). \end{aligned}$$

Lemma A8. *Let $\mathbf{W}_1, \dots, \mathbf{W}_n$ be iid uniformly distributed on S^{p-1} . Then*

$$\sqrt{\frac{p}{2n^2}} \sum_{1 \leq i \neq \ell \leq n} \langle \mathbf{W}_i, \mathbf{W}_\ell \rangle \xrightarrow{d} N(0, 1) \text{ uniformly as } n \wedge p \rightarrow \infty,$$

where $\langle \mathbf{W}_i, \mathbf{W}_\ell \rangle$ denotes the inner product of \mathbf{W}_i and \mathbf{W}_ℓ .

A.2. Proof of Theorem 1

Theorem 1 is a special case of Theorem A1 below for $m = 1$.

Theorem A1. *Let*

$$T_{n,p} = (p - 1) \ln(\sin \Theta_{m:n}) - K_{n,p},$$

where $K_{n,p}$ is defined as in (3.1). Let $G_m^*(t) = G_m(e^t), t \in \mathbb{R}$, where G_m denotes the gamma distribution with shape parameter m and scale parameter 1. Then for fixed $m = 1, 2, \dots$, $T_{n,p} \xrightarrow{d} G_m^*$ uniformly as $n \wedge p \rightarrow \infty$.

Proof. We claim that

$$T_{n_\ell, p_\ell} \xrightarrow{d} G_m^* \tag{A.5}$$

for any increasing sequences $\{n_\ell\}$ and $\{p_\ell\}$ satisfying $n_\ell \rightarrow \infty, p_\ell \rightarrow \infty$ and $(\ln n_\ell)/p_\ell \rightarrow \alpha \in [0, \infty]$ as $\ell \rightarrow \infty$. Assume for now that the claim (A.5) holds. To show that $T_{n,p} \xrightarrow{d} G_m^*$ uniformly as $n \wedge p \rightarrow \infty$, suppose to the contrary that

$\limsup_{n \wedge p \rightarrow \infty} \sup_{t \in \mathbb{R}} |\mathbb{P}(T_{n,p} \leq t) - G_m^*(t)| > 0$. Then there exist an $\varepsilon > 0$ and a sequence $\{(n_\ell, p_\ell) : \ell = 1, 2, \dots\}$ such that $\lim_{\ell \rightarrow \infty} n_\ell \wedge p_\ell = \infty$ and

$$\sup_{t \in \mathbb{R}} |\mathbb{P}(T_{n_\ell, p_\ell} \leq t) - G_m^*(t)| > \varepsilon \text{ for } \ell = 1, 2, \dots \tag{A.6}$$

There exists a subsequence $\{(n_{\ell_k}, p_{\ell_k}) : k = 1, 2, \dots\}$ such that $(\ln n_{\ell_k})/p_{\ell_k}$ converges to some value $\alpha \in [0, \infty]$. Then (A.6) contradicts (A.5), implying that $T_{n,p} \xrightarrow{d} G_m^*$ uniformly as $n \wedge p \rightarrow \infty$.

We now prove (A.5). For notational simplicity, we will deal only with the special case where $n_\ell = \ell$, $\ell = 1, 2, \dots$. The general case can be treated similarly. Specifically, we show that if $p = p_n \rightarrow \infty$ satisfies $(\ln n)/p \rightarrow \alpha \in [0, \infty]$, then $T_{n,p} = T_{n,p_n} \xrightarrow{d} G_m^*$.

Suppose $p = p_n \rightarrow \infty$ satisfies $\lim_{n \rightarrow \infty} (\ln n)/p = \alpha \in [0, \infty]$. For fixed m , since $F_p(\Theta_{m:n}) \stackrel{d}{=} U_{m:n}$, we have by Lemma A3

$$\begin{aligned} \mathbb{P}(nF_p(\Theta_{m:n}) \leq e^t) &= \mathbb{P}(nU_{m:n} \leq e^t) = \mathbb{P}\left(n \frac{S_m}{S_{n+1}} \leq e^t\right) \\ &\rightarrow \mathbb{P}(S_m \leq e^t) = G_m(e^t) = G_m^*(t). \end{aligned} \tag{A.7}$$

For fixed $t > 0$, let $t_n \in [0, 1)$ be such that

$$\frac{p-1}{2} \ln(1-t_n^2) = \min\{K_{n,p} + t, 0\}.$$

Noting that

$$K_{n,p} = K_{n,p_n} = -(\ln n)(1 + o(1)) \text{ as } n \rightarrow \infty, \tag{A.8}$$

we have for large n

$$\frac{p-1}{2} \ln(1-t_n^2) = K_{n,p} + t < 0. \tag{A.9}$$

By Lemma A2,

$$\left(1 + \frac{1}{(p-3)t_n^2}\right)^{-1} \bar{F}_p(\cos^{-1} t_n) \leq F_p(\cos^{-1} t_n) \leq \bar{F}_p(\cos^{-1} t_n),$$

implying that

$$\begin{aligned} \mathbb{P}(nF_p(\Theta_{m:n}) \leq n\bar{F}_p(\cos^{-1} t_n)) &\geq \mathbb{P}(nF_p(\Theta_{m:n}) \leq nF_p(\cos^{-1} t_n)) \\ &\geq \mathbb{P}\left(nF_p(\Theta_{m:n}) \leq \left(1 + \frac{1}{(p-3)t_n^2}\right)^{-1} n\bar{F}_p(\cos^{-1} t_n)\right). \end{aligned} \tag{A.10}$$

Recalling $\alpha = \lim_{n \rightarrow \infty} (\ln n)/p$, we claim that for every $\alpha \in [0, \infty]$, as $n \rightarrow \infty$

$$n \bar{F}_p(\cos^{-1} t_n) = e^t + o(1), \quad (\text{A.11})$$

$$p t_n^2 \rightarrow \infty, \quad (\text{A.12})$$

$$P(\cos \Theta_{m:n} \leq -t_n) \rightarrow 0. \quad (\text{A.13})$$

By (A.7), (A.10), (A.11) and (A.12),

$$P(\cos \Theta_{m:n} \geq t_n) = P(n F_p(\Theta_{m:n}) \leq n F_p(\cos^{-1} t_n)) \rightarrow G_m^*(t). \quad (\text{A.14})$$

Furthermore,

$$\begin{aligned} P(T_{n,p} \leq t) &= P\left(\frac{p-1}{2} \ln(1 - \cos^2 \Theta_{m:n}) - K_{n,p} \leq t\right) \\ &= P(\cos^2 \Theta_{m:n} \geq t_n^2) \quad (\text{by (A.9)}) \\ &= P(\cos \Theta_{m:n} \geq t_n) + P(\cos \Theta_{m:n} \leq -t_n) \\ &\rightarrow G_m^*(t) \quad (\text{by (A.13) and (A.14)}). \end{aligned}$$

It remains to establish (A.11)-(A.13). Note that by Sterling's formula (see e.g. Tricomi and Erdélyi (1951)),

$$\frac{\Gamma(p/2)}{\Gamma((p-1)/2)} = \sqrt{\frac{p}{2}} \left(1 + O\left(\frac{1}{p}\right)\right) \quad \text{as } p \rightarrow \infty. \quad (\text{A.15})$$

We have

$$\begin{aligned} \ln(n \bar{F}_p(\cos^{-1} t_n)) &= \ln \left\{ \frac{n}{\sqrt{\pi}} \frac{\Gamma(p/2)}{\Gamma((p-1)/2)} \left(\frac{(1-t_n^2)^{p-1}}{(p-1)^2 t_n^2}\right)^{1/2} \right\} \quad (\text{by (A.2)}) \\ &= \ln \left\{ n \left(\frac{(1-t_n^2)^{p-1}}{2\pi p t_n^2}\right)^{1/2} \right\} + O\left(\frac{1}{p}\right) \quad (\text{by (A.15)}) \\ &= \frac{p-1}{2} \ln(1-t_n^2) + \ln n - \frac{1}{2} \ln(p t_n^2) - \frac{1}{2} \ln(2\pi) + O\left(\frac{1}{p}\right) \\ &= K_{n,p} + t + \ln n - \frac{1}{2} \ln(p t_n^2) - \frac{1}{2} \ln(2\pi) + O\left(\frac{1}{p}\right) \quad (\text{by (A.9)}). \end{aligned} \quad (\text{A.16})$$

By (A.8) and (A.9),

$$\ln(1-t_n^2) = -\frac{2 \ln n}{p} (1 + o(1)), \quad (\text{A.17})$$

implying that

$$t_n \rightarrow (1 - e^{-2\alpha})^{1/2}, \tag{A.18}$$

where $\lim_{n \rightarrow \infty} (\ln n)/p = \alpha \in [0, \infty]$ and $e^{-\infty} := 0$.

If $\alpha = 0$, we have $t_n \rightarrow 0^+$, so that by (A.17)

$$t_n^2 = \frac{2 \ln n}{p} (1 + o(1)), \tag{A.19}$$

from which it follows that $\ln(pt_n^2) = \ln(2 \ln n) + o(1)$. By the definition of $K_{n,p}$, we have $K_{n,p} = -\ln n + (\ln \ln n)/2 + \ln(4\pi)/2 + o(1)$, so that $K_{n,p} + \ln n - \ln(pt_n^2)/2 - \ln(2\pi)/2 = o(1)$, which together with (A.16) establishes (A.11) for $\alpha = 0$. If $0 < \alpha < \infty$, we have $t_n^2 = 1 - e^{-2\alpha} + o(1)$ (by (A.18)) and $\ln(pt_n^2) = \ln \ln n - \ln \alpha + \ln(1 - e^{-2\alpha}) + o(1)$, so that $K_{n,p} + \ln n - \ln(pt_n^2)/2 - \ln(2\pi)/2 = o(1)$, which together with (A.16) establishes (A.11) for $0 < \alpha < \infty$. If $\alpha = \infty$, we have $t_n \rightarrow 1^-$, so that by the definition of $K_{n,p}$,

$$\begin{aligned} & K_{n,p} + \ln n - \frac{1}{2} \ln(pt_n^2) - \frac{1}{2} \ln(2\pi) \\ &= -\ln n + \frac{1}{2} \ln \ln n - \frac{1}{2} \ln\left(\frac{2 \ln n}{p}\right) + \frac{1}{2} \ln(4\pi) + \ln n - \frac{1}{2} \ln p - \frac{1}{2} \ln(2\pi) + o(1) \\ &= o(1), \end{aligned}$$

which together with (A.16) establishes (A.11) for $\alpha = \infty$.

Next, (A.19) holds for $\alpha = 0$, which implies (A.12). For $0 < \alpha \leq \infty$, it follows from (A.18) that $t_n \rightarrow (1 - e^{-2\alpha})^{1/2} > 0$, which implies (A.12).

Finally, to prove (A.13), note that

$$P(\cos \Theta_{m:n} \leq -t_n) \leq P\left(\Theta_{m:n} \geq \frac{\pi}{2}\right) = P\left(B\left(n, \frac{1}{2}\right) < m\right) \rightarrow 0,$$

where $B(n, 1/2)$ denotes a binomial random variable with parameters n and $1/2$ (success probability). This establishes (A.13) and completes the proof of Theorem A1.

A.3. Proofs of Theorems 2-3 and Corollary 1

We first show that if $m = m_n \rightarrow \infty$ satisfies $m/n \rightarrow 0$ and $p = p_n \rightarrow \infty$ satisfies $(\ln n)^2/p \rightarrow 0$, then

$$m \sqrt{\frac{p}{2}} \left(V_{n,p,m} - \frac{1}{m} \right) \xrightarrow{d} N(0, 1), \tag{A.20}$$

where $V_{n,p,m} = \|1/m \sum_{i=1}^m \nu_i \mathbf{V}'_i\|^2$ with $\nu_i^2 = 1 - \cos^2 \Theta_{i:n}$, and $\mathbf{V}'_1, \dots, \mathbf{V}'_m$ are i.i.d. uniformly distributed on \mathcal{S}^{p-2} , and (ν_1, \dots, ν_m) is independent of $(\mathbf{V}'_1, \dots, \mathbf{V}'_m)$.

We have

$$\begin{aligned} V_{n,p,m} &= \frac{1}{m^2} \sum_{i=1}^m \nu_i^2 \|\mathbf{V}'_i\|^2 + \frac{1}{m^2} \sum_{1 \leq i \neq \ell \leq m} \nu_i \nu_\ell \langle \mathbf{V}'_i, \mathbf{V}'_\ell \rangle \\ &= \frac{1}{m} + \frac{1}{m^2} \sum_{i=1}^m (\nu_i^2 - 1) + \frac{1}{m^2} \sum_{1 \leq i \neq \ell \leq m} \{1 + (\nu_i \nu_\ell - 1)\} \langle \mathbf{V}'_i, \mathbf{V}'_\ell \rangle \\ &= \frac{1}{m} + V'_{1,n} + V'_{2,n} + V'_{3,n}, \end{aligned} \tag{A.21}$$

where

$$\begin{aligned} V'_{1,n} &= \frac{1}{m^2} \sum_{i=1}^m (\nu_i^2 - 1) = -\frac{1}{m^2} \sum_{i=1}^m \cos^2 \Theta_{i:n}, \\ V'_{2,n} &= \frac{1}{m^2} \sum_{1 \leq i \neq \ell \leq m} \langle \mathbf{V}'_i, \mathbf{V}'_\ell \rangle, \\ V'_{3,n} &= \frac{1}{m^2} \sum_{1 \leq i \neq \ell \leq m} (\nu_i \nu_\ell - 1) \langle \mathbf{V}'_i, \mathbf{V}'_\ell \rangle. \end{aligned}$$

By Lemma A8, we have

$$m \sqrt{\frac{p}{2}} V'_{2,n} \xrightarrow{d} N(0, 1). \tag{A.22}$$

It remains to prove

$$mp^{1/2} V'_{i,n} = o_p(1), \quad i = 1, 3. \tag{A.23}$$

By Lemma A6(ii),

$$\max_{1 \leq i \leq m} \cos^2 \Theta_{i:n} = O_p \left(\frac{\ln n}{p} \right),$$

implying that $mp^{1/2} V'_{1,n} = O_p(\ln n/p^{1/2}) = o_p(1)$. To show $mp^{1/2} V'_{3,n} = o_p(1)$, note that (ν_1, \dots, ν_m) is independent of $(\mathbf{V}'_1, \dots, \mathbf{V}'_m)$ and $E[\langle \mathbf{V}'_i, \mathbf{V}'_\ell \rangle \langle \mathbf{V}'_{i'}, \mathbf{V}'_{\ell'} \rangle] = 0$ if $i \neq \ell, i' \neq \ell'$ and $\{i, \ell\} \neq \{i', \ell'\}$. Also, for $i \neq \ell, E\langle \mathbf{V}'_i, \mathbf{V}'_\ell \rangle^2 = \int_0^\pi \cos^2(\theta) dF_{p-1}(\theta) = 1/(p-1)$, where F_p is defined as in (A.1). We have

$$E V'_{3,n}{}^2 = \frac{2}{m^4} \sum_{1 \leq i \neq \ell \leq m} E[(\nu_i \nu_\ell - 1)^2] E\langle \mathbf{V}'_i, \mathbf{V}'_\ell \rangle^2$$

$$\begin{aligned}
 &= \frac{2}{m^4} \sum_{1 \leq i \neq \ell \leq m} \mathbb{E}[(\nu_i \nu_\ell - 1)^2] \frac{1}{p-1} \\
 &= o\left(\frac{1}{m^2 p}\right),
 \end{aligned} \tag{A.24}$$

since $|\nu_i| \leq 1$ and $\nu_i \nu_\ell - 1 \rightarrow 0$ in probability uniformly in $1 \leq i \neq \ell \leq m$. It follows from (A.24) that $mp^{1/2}V'_{3,n} = o_p(1)$. This proves (A.23) and completes the proof of (A.20).

Proof of Theorem 2. Since by (A.3) $\rho_{n,p,m}^2 \stackrel{d}{=} A_{n,p,m}/(A_{n,p,m} + V_{n,p,m})$, we have

$$\begin{aligned}
 &\rho_{n,p,m}^2 - \frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}} \\
 &\stackrel{d}{=} \frac{A_{n,p,m} - \beta_{n,p,m}/m}{(A_{n,p,m} + V_{n,p,m})(1 + \beta_{n,p,m})} + \frac{(1/m - V_{n,p,m})\beta_{n,p,m}}{(A_{n,p,m} + V_{n,p,m})(1 + \beta_{n,p,m})}.
 \end{aligned} \tag{A.25}$$

Since $\beta_{n,p,m} = (m/p) \{2 \ln(n/m) - \ln \ln(n/m) - \ln(4\pi) + 2\}$, it follows from Lemma A7(i) and (A.20) that

$$\begin{aligned}
 p(A_{n,p,m} - \frac{1}{m}\beta_{n,p,m}) &= O_p(1), \\
 mV_{n,p,m} &= 1 + o_p(1), \\
 p\beta_{n,p,m}V_{n,p,m} &= (2 + o_p(1)) \ln\left(\frac{n}{m}\right).
 \end{aligned}$$

Thus,

$$\begin{aligned}
 \frac{A_{n,p,m} - \beta_{n,p,m}/m}{(A_{n,p,m} + V_{n,p,m})(1 + \beta_{n,p,m})} &= \frac{p(A_{n,p,m} - \beta_{n,p,m}/m)}{(p\beta_{n,p,m}A_{n,p,m} + p\beta_{n,p,m}V_{n,p,m})} \frac{\beta_{n,p,m}}{(1 + \beta_{n,p,m})} \\
 &= o_p(1) \frac{\beta_{n,p,m}}{(1 + \beta_{n,p,m})}, \\
 \frac{(1/m - V_{n,p,m})\beta_{n,p,m}}{(A_{n,p,m} + V_{n,p,m})(1 + \beta_{n,p,m})} &= \frac{(1 - mV_{n,p,m})}{(mA_{n,p,m} + mV_{n,p,m})} \frac{\beta_{n,p,m}}{(1 + \beta_{n,p,m})} \\
 &= o_p(1) \frac{\beta_{n,p,m}}{(1 + \beta_{n,p,m})}.
 \end{aligned}$$

We have by (A.25),

$$\rho_{n,p,m}^2 = \frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}}(1 + o_p(1)).$$

The proof is complete.

Proof of Theorem 3. By (A.21)-(A.23),

$$\begin{aligned} m\sqrt{\frac{p}{2}}\left(V_{n,p,m} - \frac{1}{m}\right) &= m\sqrt{\frac{p}{2}}(V'_{1,n} + V'_{2,n} + V'_{3,n}) \\ &= m\sqrt{\frac{p}{2}}V'_{2,n} + o_p(1). \end{aligned} \quad (\text{A.26})$$

Let

$$\begin{aligned} Z_{1,n} &= p\sqrt{\frac{m}{8}}\left(A_{n,p,m} - \frac{1}{m}\beta_{n,p,m} + \frac{2}{p^2}\left(\ln\frac{n}{m}\right)^2\right), \\ Z_{2,n} &= m\sqrt{\frac{p}{2}}V'_{2,n}, \\ \gamma_n &= (A_{n,p,m} + V_{n,p,m})(1 + \beta_{n,p,m}). \end{aligned}$$

We have by (A.25) and (A.26)

$$\begin{aligned} &\rho_{n,p,m}^2 - \frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}} \\ &\stackrel{d}{=} \gamma_n^{-1} \left\{ \frac{1}{p\sqrt{m/8}}Z_{1,n} - \frac{\beta_{n,p,m}}{m\sqrt{p/2}}m\sqrt{\frac{p}{2}}\left(V_{n,p,m} - \frac{1}{m}\right) \right\} - \gamma_n^{-1} \frac{2}{p^2} \left(\ln\frac{n}{m}\right)^2 \\ &= \gamma_n^{-1} \left\{ \sqrt{\frac{8}{mp^2}}Z_{1,n} - \sqrt{\frac{2}{m^2p}}\beta_{n,p,m}(Z_{2,n} + o_p(1)) \right\} - \gamma_n^{-1} \frac{2}{p^2} \left(\ln\frac{n}{m}\right)^2 \\ &= \gamma_n^{-1} \left(\frac{8}{mp^2} + \frac{2}{m^2p}\beta_{n,p,m}^2 \right)^{1/2} \{c_{1,n}Z_{1,n} + c_{2,n}(Z_{2,n} + o_p(1))\} \\ &\quad - \gamma_n^{-1} \frac{2}{p^2} \left(\ln\frac{n}{m}\right)^2, \end{aligned} \quad (\text{A.27})$$

where

$$\begin{aligned} c_{1,n} &= \sqrt{\frac{8}{mp^2}} \left(\frac{8}{mp^2} + \frac{2}{m^2p}\beta_{n,p,m}^2 \right)^{-1/2}, \\ c_{2,n} &= -\sqrt{\frac{2}{m^2p}}\beta_{n,p,m} \left(\frac{8}{mp^2} + \frac{2}{m^2p}\beta_{n,p,m}^2 \right)^{-1/2}. \end{aligned}$$

Since $\rho_{n,p,m}^2 \stackrel{d}{=} A_{n,p,m}/(A_{n,p,m} + V_{n,p,m})$, we have by Theorem 2

$$\frac{A_{n,p,m}}{A_{n,p,m} + V_{n,p,m}} = \frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}}(1 + o_p(1)). \quad (\text{A.28})$$

It follows from Lemma A7(i) and $(p/m)\beta_{n,p,m} = 2 \ln(n/m)(1 + o(1))$ that

$$\frac{mA_{n,p,m}}{\beta_{n,p,m}} = \frac{pA_{n,p,m}}{(p/m)\beta_{n,p,m}} = 1 + o_p(1). \tag{A.29}$$

So we have

$$\begin{aligned} & \gamma_n \left(\frac{8}{mp^2} + \frac{2}{m^2p} \beta_{n,p,m}^2 \right)^{-1/2} \\ &= \frac{pm}{\sqrt{8m + 2p\beta_{n,p,m}^2}} (A_{n,p,m} + V_{n,p,m})(1 + \beta_{n,p,m}) \\ &= \frac{pmA_{n,p,m}}{\sqrt{8m + 2p\beta_{n,p,m}^2}} \frac{A_{n,p,m} + V_{n,p,m}}{A_{n,p,m}} (1 + \beta_{n,p,m}) \\ &= \frac{pmA_{n,p,m}/\beta_{n,p,m}}{\sqrt{8m + 2p\beta_{n,p,m}^2}} (1 + \beta_{n,p,m})^2 (1 + o_p(1)) \quad (\text{by (A.28)}) \\ &= \frac{p}{\sqrt{8m + 2p\beta_{n,p,m}^2}} (1 + \beta_{n,p,m})^2 (1 + o_p(1)) \quad (\text{by (A.29)}) \\ &= \alpha_{n,p,m} (1 + o_p(1)), \end{aligned} \tag{A.30}$$

where $\alpha_{n,p,m} = p (8m + 2p\beta_{n,p,m}^2)^{-1/2} (1 + \beta_{n,p,m})^2$.

Also,

$$\begin{aligned} 0 &< \frac{2}{p^2} \left(\ln \frac{n}{m} \right)^2 \left(\frac{8}{mp^2} + \frac{2}{m^2p} \beta_{n,p,m}^2 \right)^{-1/2} \\ &\leq \frac{2}{p^2} \left(\ln \frac{n}{m} \right)^2 \left(\frac{2}{m^2p} \beta_{n,p,m}^2 \right)^{-1/2} \\ &= \frac{2}{p^2} \left(\ln \frac{n}{m} \right)^2 \left\{ \frac{2}{p^3} \left(\frac{p}{m} \beta_{n,p,m} \right)^2 \right\}^{-1/2} \\ &= \sqrt{\frac{2}{p}} \left(\ln \frac{n}{m} \right)^2 \left(2 \ln \frac{n}{m} (1 + o(1)) \right)^{-1} \\ &= \frac{1}{\sqrt{2p}} \ln \frac{n}{m} (1 + o(1)) = o(1), \end{aligned}$$

which together with (A.30) implies that

$$\begin{aligned} & \frac{\alpha_{n,p,m}}{\gamma_n} \frac{2}{p^2} \left(\ln \frac{n}{m} \right)^2 \\ &= \left\{ \frac{\alpha_{n,p,m}}{\gamma_n} \left(\frac{8}{mp^2} + \frac{2}{m^2p} \beta_{n,p,m}^2 \right)^{1/2} \right\} \left\{ \frac{2}{p^2} \left(\ln \frac{n}{m} \right)^2 \left(\frac{8}{mp^2} + \frac{2}{m^2p} \beta_{n,p,m}^2 \right)^{-1/2} \right\} \\ &= (1 + o_p(1))o(1) = o_p(1). \end{aligned} \tag{A.31}$$

It follows from (A.27), (A.30), and (A.31) that

$$\begin{aligned} & \alpha_{n,p,m} \left(\rho_{n,p,m}^2 - \frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}} \right) \\ & \stackrel{d}{=} \frac{\alpha_{n,p,m}}{\gamma_n} \left(\frac{8}{mp^2} + \frac{2}{m^2p} \beta_{n,p,m}^2 \right)^{1/2} \{c_{1,n}Z_{1,n} + c_{2,n}Z_{2,n}(1 + o_p(1))\} \\ & \quad - \frac{\alpha_{n,p,m}}{\gamma_n} \frac{2}{p^2} \left(\ln \frac{n}{m} \right)^2 \\ & = (1 + o_p(1)) \{c_{1,n}Z_{1,n} + c_{2,n}Z_{2,n}(1 + o_p(1))\} + o_p(1). \end{aligned} \tag{A.32}$$

Note that $c_{1,n}$ and $c_{2,n}$ are constants (depending on n, p_n, m_n), which satisfy $c_{1,n}^2 + c_{2,n}^2 = 1$. By Lemma A7(iii),

$$\begin{aligned} -Z_{1,n} &= \sqrt{\frac{m}{8}} \left\{ -p A_{n,p,m} + 2 \ln \frac{n}{m} - \ln \ln \frac{n}{m} - \ln(4\pi) + 2 - \frac{2}{p} \left(\ln \frac{n}{m} \right)^2 \right\} \\ &\xrightarrow{d} N(0, 1). \end{aligned}$$

By (A.22), $Z_{2,n} \xrightarrow{d} N(0, 1)$. Note that $Z_{1,n}$ and $Z_{2,n}$ are independent (since $A_{n,p,m}$ and $V'_{2,n}$ are independent). We have

$$c_{1,n}Z_{1,n} + c_{2,n}Z_{2,n} \xrightarrow{d} N(0, 1),$$

which together with (A.32) implies that

$$\alpha_{n,p,m} \left(\rho_{n,p,m}^2 - \frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}} \right) \xrightarrow{d} N(0, 1).$$

The proof is complete.

Proof of Corollary 1. Part (i) follows immediately from Theorem 2. To prove part (ii), we have by part (i) and Theorem 3 that

$$\begin{aligned} & 2\alpha_{n,p,m} \sqrt{\frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}}} \left(\rho_{n,p,m} - \sqrt{\frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}}} \right) \\ & = \frac{2\sqrt{\beta_{n,p,m}/(1 + \beta_{n,p,m})}}{\rho_{n,p,m} + \sqrt{\beta_{n,p,m}/(1 + \beta_{n,p,m})}} \alpha_{n,p,m} \left(\rho_{n,p,m}^2 - \frac{\beta_{n,p,m}}{1 + \beta_{n,p,m}} \right) \\ & \xrightarrow{d} N(0, 1), \end{aligned}$$

completing the proof.

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