# STATISTICAL-PHYSICAL ESTIMATION OF POLLUTION EMISSION 

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## Supplementary Material

## S1 Generation of wind field data

The wind velocity in the $s_{1}$-direction is

$$
\begin{equation*}
u_{1}(\boldsymbol{s}, t)=\operatorname{Real}\left(\sum_{l=1}^{4} \sum_{m=1}^{4}\left(\widehat{u}_{r}(l, m, t)+i \widehat{u}_{i}(l, m, t)\right) e^{i \boldsymbol{k}(l, m) \cdot \boldsymbol{s}}\right), \tag{S1.1}
\end{equation*}
$$

in which $\boldsymbol{k}(l, m)=\left(\frac{2 \pi}{40 \mathrm{~km}} l, \frac{2 \pi}{40 \mathrm{~km}} m\right)$ is the wavenumber. The Fourier coefficients are computed by solving the following stochastic differential equation;

$$
\begin{equation*}
\delta \widehat{u}_{r, i}(l, m, t)=-\frac{\widehat{u}_{r, i}(l, m, t)}{T_{L}} \delta t+S(l, m) \delta W \tag{S1.2}
\end{equation*}
$$

in which $T_{L}$ is a relaxation timescale, S is a scale parameter, and W denotes the Wiener process. Note that the solution of the stochastic differential equation is given by the Ornstein-Uhlenbeck process. In this study, $T_{L}=6$
hours and $S=2 \sqrt{\frac{2}{T_{L}\left(l^{2}+m^{2}\right)}}$ are used. The velocity in the $s_{2}$-direction is computed from the mass conservation constraint;

$$
\begin{equation*}
\frac{\partial u_{1}(\boldsymbol{s}, t)}{\partial s_{1}}+\frac{\partial u_{2}(\boldsymbol{s}, t)}{\partial s_{2}}=0 . \tag{S1.3}
\end{equation*}
$$

In other words,

$$
\begin{equation*}
u_{2}(\boldsymbol{s}, t)=-\operatorname{Real}\left(\sum_{l=1}^{p} \sum_{m=1}^{p} \frac{L_{2}}{L_{1}} \frac{l}{m}\left(\widehat{u}_{r}(l, m, t)+i \widehat{u}_{i}(l, m, t)\right) e^{i \boldsymbol{k}(l, m) \cdot \boldsymbol{s}}\right) . \tag{S1.4}
\end{equation*}
$$

The diffusivity matrix $\boldsymbol{K}(\boldsymbol{s}, t ; \boldsymbol{u})$ is computed by an isotropic Smagorinsky model, which is typically used in the atmospheric dynamics models (Byun and Schere, 2006);

$$
\begin{aligned}
& K_{i j}(\boldsymbol{s}, t)=K_{h}(\boldsymbol{s}, t) \delta_{i j}, \\
& K_{h}(\boldsymbol{s}, t)=\left(C_{s} \Delta\right)^{2} \sqrt{\left(\frac{\partial u_{1}(\boldsymbol{s}, t)}{\partial s_{1}}-\frac{\partial u_{2}(\boldsymbol{s}, t)}{\partial s_{2}}\right)+\left(\frac{\partial u_{1}(\boldsymbol{s}, t)}{\partial s_{2}}+\frac{\partial u_{2}(\boldsymbol{s}, t)}{\partial s_{1}}\right)} .
\end{aligned}
$$

Here, $\delta_{i j}$ is the Kronecker delta, $C_{s}(=0.1)$ is the Smagorinsky coefficient, and the length scale $\Delta=40 \mathrm{~km} / 2 \pi \times 4$.

## S2 Derivation of ADMM Algorithm

We define new variables, $\boldsymbol{\beta}_{\mathrm{mse}}, \boldsymbol{\beta}_{\mathrm{gl}}, \boldsymbol{\beta}_{\mathrm{nuc}}, \boldsymbol{\beta}_{\mathrm{nn}}$ and rewrite the main objective function:

$$
\begin{array}{rc}
\operatorname{minimize} & f_{\mathrm{mse}}\left(\boldsymbol{\beta}_{\mathrm{mse}}\right)+f_{\mathrm{gl}}\left(\boldsymbol{\beta}_{\mathrm{gl}}\right)+f_{\mathrm{nuc}}\left(\boldsymbol{\beta}_{\mathrm{nuc}}\right)+f_{n n}\left(\boldsymbol{\beta}_{\mathrm{nn}}\right) \\
\text { subject to } & \boldsymbol{\beta}_{\mathrm{mse}}=\boldsymbol{\beta}_{\mathrm{gl}}=\boldsymbol{\beta}_{\mathrm{nuc}}=\boldsymbol{\beta}_{\mathrm{nn}} \tag{S2.6}
\end{array}
$$

We start from a feasible solution, $\boldsymbol{\beta}_{\mathrm{mse}}=\boldsymbol{\beta}_{\mathrm{gl}}=\boldsymbol{\beta}_{\mathrm{nuc}}=\boldsymbol{\beta}_{\mathrm{nn}}=0$. At each stage of the algorithm, we separately minimize the components of (S2.5) using proximal methods. Let $\boldsymbol{\beta}_{\mathrm{mse}}^{(m)}$ denote the solution for minimizing $f_{\text {mse }}$ at $m^{\text {th }}$ iteration; we similarly define $\boldsymbol{\beta}_{g}^{(m)}$ for $g=\{\mathrm{gl}$, nuc, $\operatorname{pos}\}$. Once $m^{\text {th }}$ iteration is complete, the average of each update can be defined to be

$$
\overline{\boldsymbol{\beta}}^{m}=\left(\boldsymbol{\beta}_{\mathrm{mse}}^{(m)}+\boldsymbol{\beta}_{\mathrm{gl}}^{(m)}+\boldsymbol{\beta}_{\mathrm{nuc}}^{(m)}+\boldsymbol{\beta}_{\mathrm{nn}}^{(m)}\right) / 4 .
$$

The dual variables, $\boldsymbol{u}_{g}^{(m)}$ for $g=\{\mathrm{mse}, \mathrm{gl}$, nuc, nn$\}$ give the deviation of $\boldsymbol{\beta}_{g}^{(m)}$ from the average. The algorithm merges $\boldsymbol{\beta}_{g}^{(m)}$ toward $\overline{\boldsymbol{\beta}}^{(m)}$ by making $\boldsymbol{u}_{g}^{(m)}$ close to zero.

At the $m^{\text {th }}$ step the algorithm first minimizes the separate functions. For $g=\{\mathrm{mse}, \mathrm{gl}$, nuc $\}$, we set

$$
\begin{equation*}
\boldsymbol{\beta}_{g}^{(m+1)} \leftarrow \arg \min _{\boldsymbol{\beta}}\left(f_{g}(\boldsymbol{\beta})+\left(\boldsymbol{u}_{g}^{(m)}\right)^{\top}\left(\boldsymbol{\beta}-\overline{\boldsymbol{\beta}}^{(m)}\right)+\frac{1}{2 \rho}\left\|\boldsymbol{\beta}-\overline{\boldsymbol{\beta}}^{(m)}\right\|_{F}^{2}\right), \tag{S2.7}
\end{equation*}
$$

where $\rho$ is a step size for the algorithm. The minimization over $f_{n n}$ is done via projection:

$$
\begin{equation*}
\boldsymbol{\beta}_{\mathrm{nn}}^{(m+1)} \leftarrow \max \left\{0, \overline{\boldsymbol{\beta}}^{(m)}-\boldsymbol{u}_{\mathrm{nn}}^{(m)}\right\}, \tag{S2.8}
\end{equation*}
$$

where max refers to the component-wise maximum.
Finally, we compute the average consensus variable and update the dual variables:

$$
\begin{array}{ll}
\overline{\boldsymbol{\beta}}^{(m+1)} \leftarrow\left(\boldsymbol{\beta}_{\mathrm{mse}}^{(m+1)}+\boldsymbol{\beta}_{\mathrm{gl}}^{(m+1)}+\boldsymbol{\beta}_{\mathrm{nuc}}^{(m+1)}+\boldsymbol{\beta}_{\mathrm{nn}}^{(m+1)}\right) / 4, \\
\boldsymbol{u}_{g}^{(m+1)} \leftarrow \boldsymbol{u}_{g}^{(m)}+\left(\boldsymbol{\beta}_{g}^{(m+1)}-\overline{\boldsymbol{\beta}}^{(m+1)}\right) \quad \text { for } g=\{\mathrm{mse}, \mathrm{gl}, \mathrm{nuc}, \mathrm{nn}\} . \tag{S2.9}
\end{array}
$$

Now, we analyze the solutions of equation (S2.7) for each component in detail.

The first function, $f_{\text {mse }}$, is for updating $\boldsymbol{\beta}$ with respect to the sum of prediction errors. For $k=1, \ldots, 24$, and $i=1, \ldots, n$, we define the matrices $\mathbf{X}(i, k) \in \mathbb{R}^{(T / 24) \times p}$ as the emissions to location $i$ at hour $k$, that is

$$
\mathbf{X}_{t, j}(i, k)=\left\{X_{t, i j}:(t \bmod 24)=k-1\right\} .
$$

We then combine these matrices row-wise and define $\mathbf{X}(k)$ for $k=1, \ldots, 24$ as

$$
\mathbf{X}(k)^{\top}=\left[\mathbf{X}(1, k)^{\top}, \ldots, \mathbf{X}(n, k)^{\top}\right] .
$$

Similarly, we define $\mathbf{y}(i, k) \in \mathbb{R}^{T / 24 \times 1}$ as the pollution levels at sensor $i$ for hour $k$. We then set, $\mathbf{y}(k)^{\top}=\left[\mathbf{y}(1, k)^{\top}, \ldots, \mathbf{y}(n, k)^{\top}\right]$, i.e. all
observations from all sensors for hour $k$. Then, according to (S2.7),

$$
\boldsymbol{\beta}_{\mathrm{mse}}^{(m+1)}=\arg \min _{\boldsymbol{\beta}}\left(\sum_{k=1}^{24}\left\|\mathbf{y}(k)-\mathbf{X}(k) \boldsymbol{\beta}_{:, k}\right\|_{2}^{2}+\left(\boldsymbol{u}_{\mathrm{mse}}^{m}\right)^{\top}\left(\boldsymbol{\beta}-\overline{\boldsymbol{\beta}}^{(m)}\right)+\frac{1}{2 \rho}\left\|\boldsymbol{\beta}-\overline{\boldsymbol{\beta}}^{(m)}\right\|_{F}^{2}\right)
$$

where $\boldsymbol{\beta}_{:, k}$ is the $k^{\text {th }}$ column of $\boldsymbol{\beta}$. Note that the problem is separable over $k$, and differentiating with respect to $\boldsymbol{\beta}_{:, k}$ reduces to solving normal equations given by

$$
\left(\mathbf{X}(k)^{\top} \mathbf{X}(k)+\frac{1}{2 \rho} \mathbf{I}_{p \times p}\right) \boldsymbol{\beta}_{:, k}=\mathbf{X}(k)^{\top} \mathbf{y}(k)+\frac{1}{2 \rho}\left(\overline{\boldsymbol{\beta}}_{:, k}^{(m)}-\rho \boldsymbol{u}_{\mathrm{mse}}^{(m)}(k)\right),
$$

where with an abuse of notation $\boldsymbol{u}_{\mathrm{mse}}^{(m)}(k)$ refers to $k^{\text {th }}$ column of $\boldsymbol{u}_{\mathrm{mse}}^{(m)}$.
The proximal step for $\boldsymbol{\beta}_{\mathrm{gl}}$ is given by a soft-thresholding shrinkage operator for the group lasso. That is, $\boldsymbol{\beta}_{\mathrm{gl}}^{(m+1)}$ is obtained by

$$
\boldsymbol{\beta}_{\mathrm{gl}}^{(m+1)} \leftarrow \arg \min _{\boldsymbol{\beta}}\left(\lambda_{\mathrm{gl} 1}\|\boldsymbol{\beta}\|_{2}+\boldsymbol{u}_{\mathrm{gl}}^{(m)}\left(\boldsymbol{\beta}-\overline{\boldsymbol{\beta}}^{(m)}\right)+\frac{1}{2 \rho}\left\|\boldsymbol{\beta}-\overline{\boldsymbol{\beta}}^{(m)}\right\|_{F}^{2}\right)
$$

Note that, the Karush-Kuhn-Tucker (KKT) conditions for optimality are given by,

$$
0 \subseteq \frac{\lambda_{\mathrm{gl}}}{\|\boldsymbol{\beta}\|_{2}}+\boldsymbol{u}_{\mathrm{gl}}^{(m)}+\frac{1}{\rho}\left(\boldsymbol{\beta}-\overline{\boldsymbol{\beta}}^{(m)}\right)
$$

which gives
$\boldsymbol{\beta}_{\mathrm{gl}}^{(m+1)}= \begin{cases}\operatorname{Sign}\left(\overline{\boldsymbol{\beta}}^{(m)}-\rho \boldsymbol{u}_{\mathrm{gl}}^{(m)}\right)\left(1-\frac{\lambda_{\mathrm{g} 1} \rho}{\left\|\overline{\boldsymbol{\beta}}^{(m)}-\rho \boldsymbol{u}_{\mathrm{g} 1}^{(m)}\right\|_{2}}\right)\left(\overline{\boldsymbol{\beta}}^{(m)}-\rho \boldsymbol{u}_{\mathrm{gl}}^{(m)}\right) & \text { if }\left\|\overline{\boldsymbol{\beta}}^{(m)}-\boldsymbol{u}_{\mathrm{gl}}^{(m)}\right\|_{2} \geq \lambda_{\mathrm{gl}} \rho, \\ 0 & \text { otherwise. }\end{cases}$
In other words, this step applies a hard-thresholding operator on the groups,
which are given by pollution sites.

For the nuclear norm regularization, $f_{\text {nuc }}$, consider the SVD of $\boldsymbol{\beta}^{(m)}-$ $\boldsymbol{u}_{\text {nuc }}^{(m)}=\boldsymbol{U}^{(m)} \boldsymbol{\Sigma}^{(m)} \boldsymbol{V}^{(m)^{\top}}$. Then, $\boldsymbol{\beta}_{\text {nuc }}^{(m+1)}$ is given by (Theorem 2.1, Cai et al. 2010

$$
\boldsymbol{\beta}_{\mathrm{nuc}}^{(m+1)}=\boldsymbol{U}^{(m)} \tilde{\boldsymbol{\Sigma}}^{(m)} \boldsymbol{V}^{(m)^{\top}}
$$

where $\tilde{\boldsymbol{\Sigma}}^{(m)}=\left(\boldsymbol{\Sigma}^{(m)}-\lambda_{\text {nuc }} \rho \mathbb{I}_{p \times p}\right)_{+}$, and $(\cdot)_{+}$is applied element-wise with $(z)_{+}=\max (0, z)$.

## S3 Configurations for Nuclear Norm "Toy Example"

 SimulationFor the toy example simulation in Section 3, we generate observations from the multivariate linear regression equation,

$$
\mathbf{y}=\mathbf{X} \boldsymbol{\beta}+\epsilon
$$

where $\mathbf{y} \in \mathbb{R}^{T \times 24}, \mathbf{X} \in \mathbb{R}^{T \times p}$ and $\boldsymbol{\beta} \in \mathbb{R}^{p \times 24}$ are the response, predictor and coefficient matrices, respectively.

The true coefficient matrix, $\boldsymbol{\beta}$, contains three linearly independent rows, which are given by $\beta_{\text {Type } 1}, \beta_{\text {Type } 2}, \beta_{\text {Type } 3} \in \mathbb{R}^{1 \times 24}$. Rest of the rows of $\boldsymbol{\beta}$ are given by copies of these vectors. We repeat each vector four times; hence $\boldsymbol{\beta}$ contains 12 rows and has matrix rank 3.

The contents of $\beta_{\text {Type1 }}$ and other types for $j=1, \ldots, 24$ are given by:

$$
\begin{aligned}
& \beta_{\text {Type } 1, \mathrm{j}}=1 \\
& \beta_{\mathrm{Type} 2, j}=\frac{1.5}{144}\left(24(j-1)-(j-1)^{2}\right) \\
& \beta_{\mathrm{Type} 3, j}= \begin{cases}2 & \text { if } j \in\{7,8,9,16,17,18\} \\
0 & \text { otherwise }\end{cases}
\end{aligned}
$$

We fill the entries of $\mathbf{X}$ by drawing independent copies from a Uniform $[0,1]$ distribution. The error matrix, $\epsilon$ is similarly filled by independent copies of a standard normal distribution.

## Bibliography

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